

# IMAGE RESTORATION

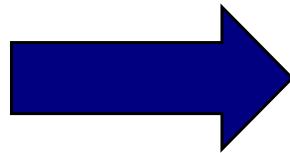
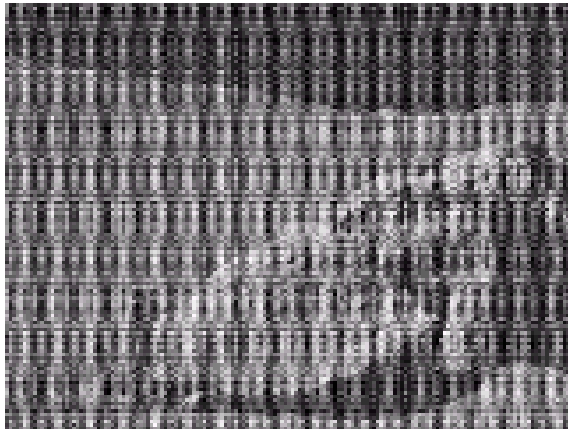


# Outline

- A model of the image degradation / restoration process
- Noise models
- Restoration in the presence of noise only – spatial filtering
- Inverse filtering & Wiener filtering
- Constrained Least square filtering
- Geometric mean filter
- Geometric and spatial transformation

# What is Image Restoration?

- Image restoration is to restore a degraded image back to the original image
- Image enhancement is to manipulate the image so that it is suitable for a specific application.



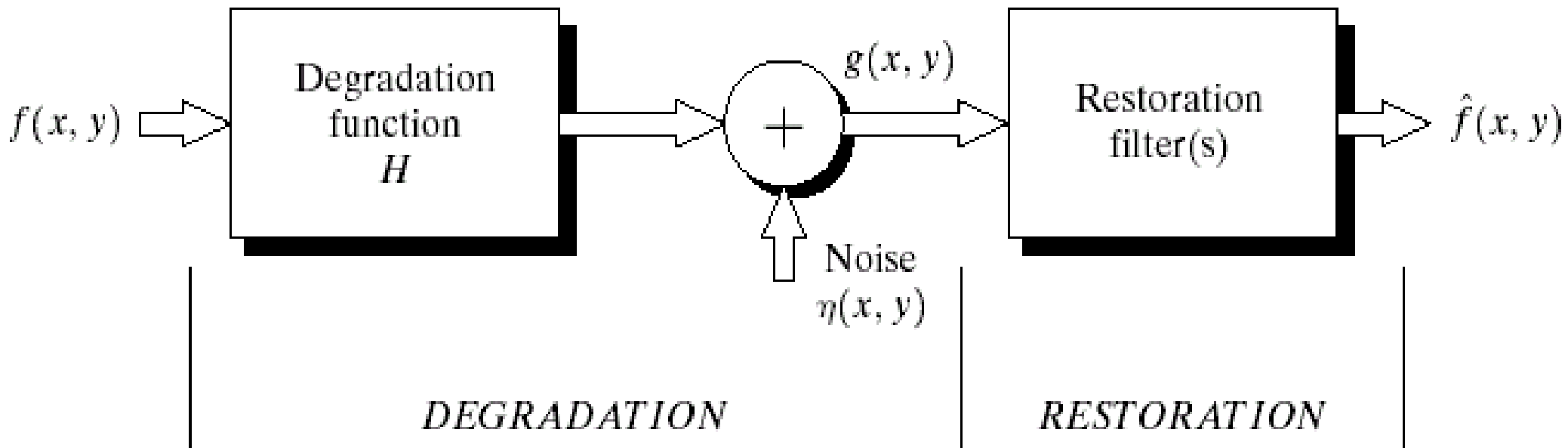
# Image Restoration

Image restoration attempts to restore images that have been degraded

- Identify the degradation process and attempt to reverse it
- Similar to image enhancement, but more objective



# A model of the image degradation/restoration process



$$g(x,y)=f(x,y)*h(x,y)+\eta(x,y) \quad \text{– Spatial domain}$$

$$G(u,v)=F(u,v)H(u,v)+N(u,v) \quad \text{– Frequency domain}$$

# A model of the image degradation/ restoration process

- Where,

$f(x,y)$  - input image

$f^{\wedge}(x,y)$  - estimated original image

$g(x,y)$  - degraded image

$h(x,y)$  - degradation function

$\eta(x,y)$  - additive noise term

# Noise models

- Sources of noise
  - Image acquisition (digitization) - Imaging sensors can be affected by ambient conditions
  - Image transmission - Interference can be added to an image during transmission
- Spatial properties of noise
  - Statistical behavior of the gray-level values of pixels
  - Noise parameters, correlation with the image
- Frequency properties of noise
  - Fourier spectrum
  - Ex. white noise (a constant Fourier spectrum)



# Gaussian noise

- Mathematical tractability in spatial and frequency domains
- Used frequently in practice
- Electronic circuit noise and sensor noise

$$p(z) = \frac{1}{\sqrt{2\pi\sigma}} e^{-\frac{(z-\mu)^2}{2\sigma^2}}$$

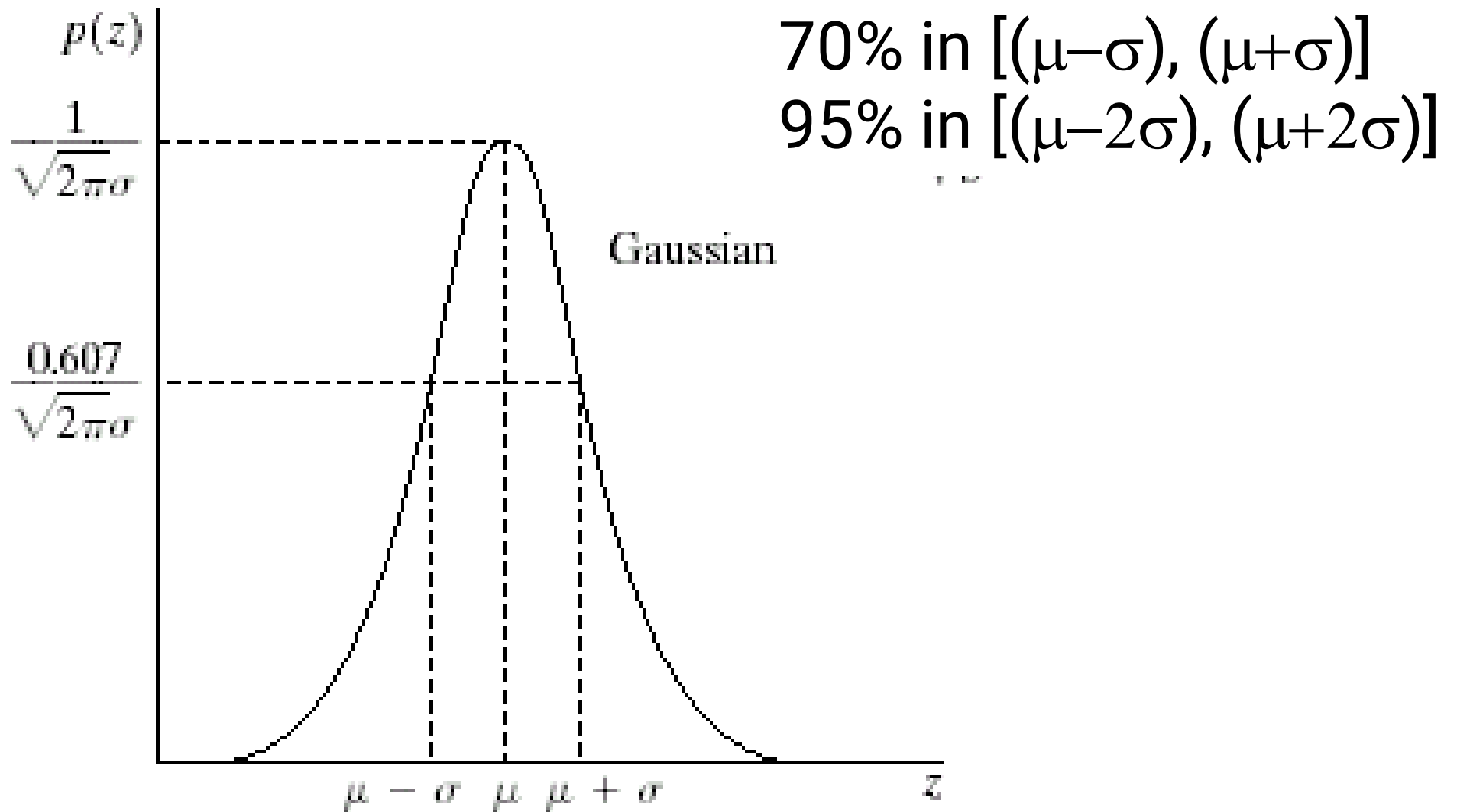
Intensity      mean      variance

Note:  $\int_{-\infty}^{\infty} p(z) dz = 1$





# Gaussian noise PDF



# Rayleigh noise

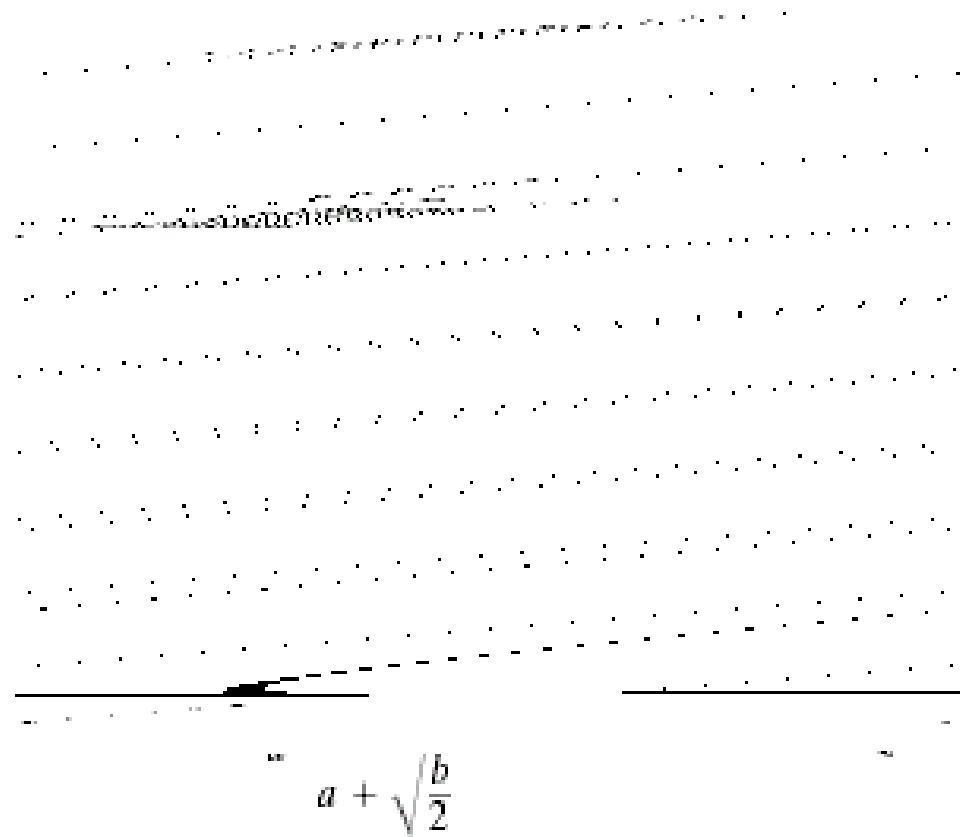
$$p(z) = \begin{cases} \frac{2}{b} (z - a) e^{-(z-a)^2 / b} & \text{for } z \geq a \\ 0 & \text{for } z < a \end{cases}$$

– The mean and variance of this density are given by

$$\mu = a + \sqrt{\pi b / 4} \quad \text{and} \quad \sigma^2 = \frac{b(4 - \pi)}{4}$$

– a and b can be obtained through mean and variance

# Rayleigh noise PDF



# Erlang (Gamma) noise

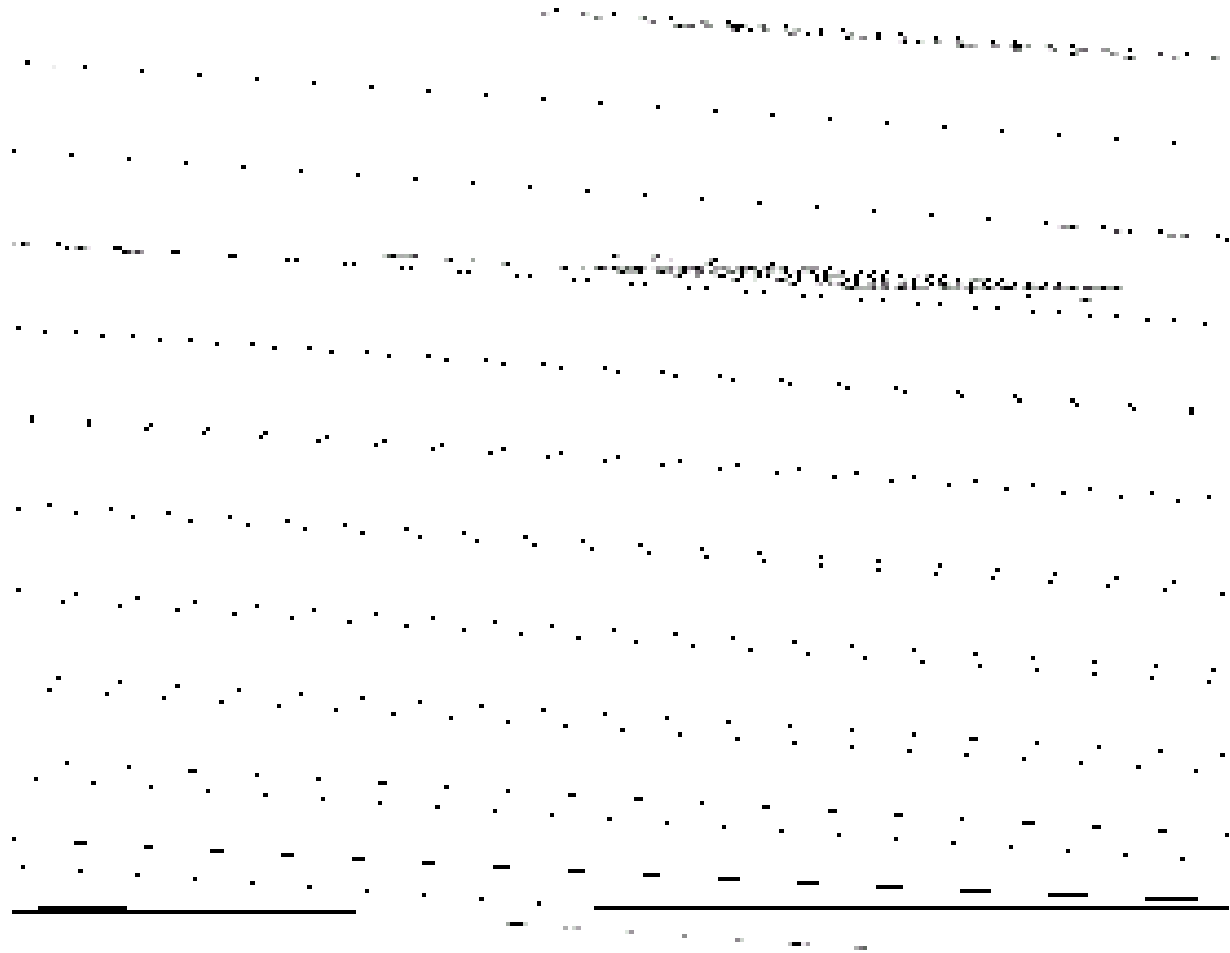
$$p(z) = \begin{cases} \frac{a^b z^{b-1}}{(b-1)!} e^{-az} & \text{for } z \geq 0 \\ 0 & \text{for } z < 0 \end{cases}$$

- The mean and variance of this density are given by

$$\mu = b/a \quad \text{and} \quad \sigma^2 = \frac{b}{a^2}$$

- a and b can be obtained through mean and variance

# Gamma noise (PDF)



# Exponential noise

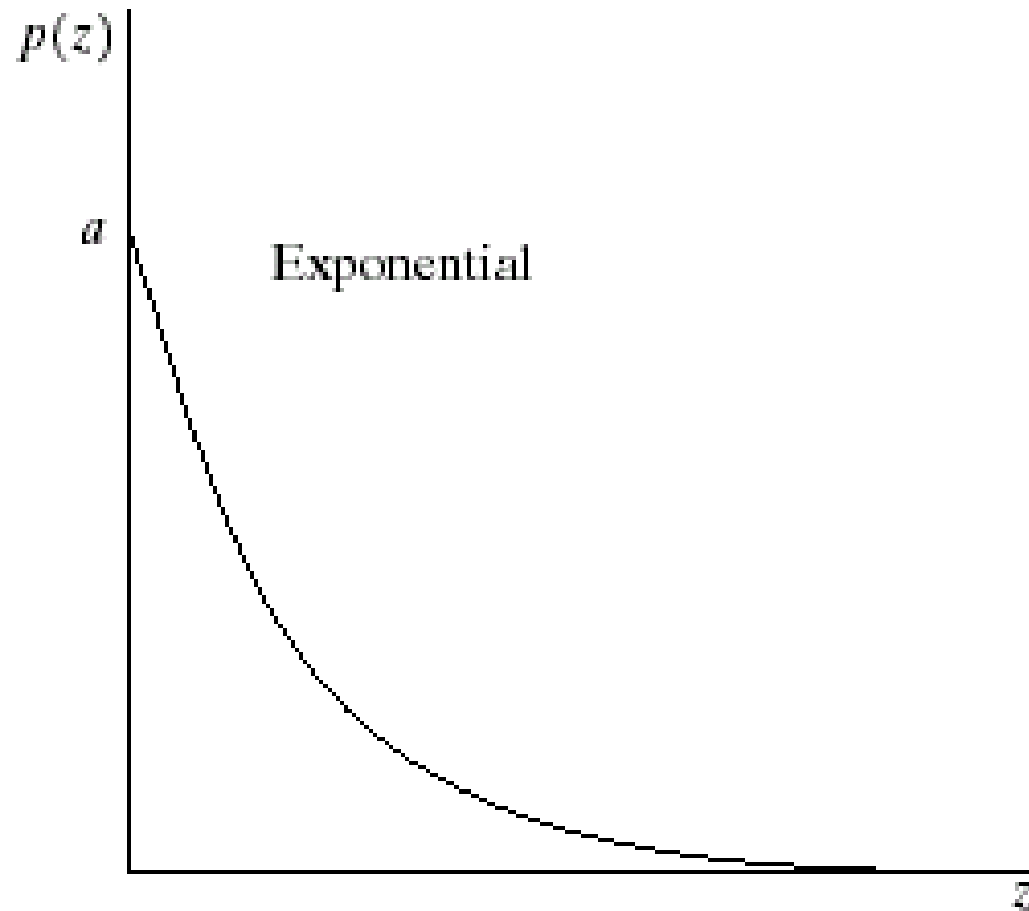
$$p(z) = \begin{cases} ae^{-az} & \text{for } z \geq 0 \\ 0 & \text{for } z < 0 \end{cases}$$

The mean and variance of this density are given by

$$\mu = 1/a \quad \text{and} \quad \sigma^2 = \frac{1}{a^2}$$

# Exponential Noise PDF

Special case of Erlang PDF with  $b=1$



# Uniform noise

- Less practical, used for random number generator

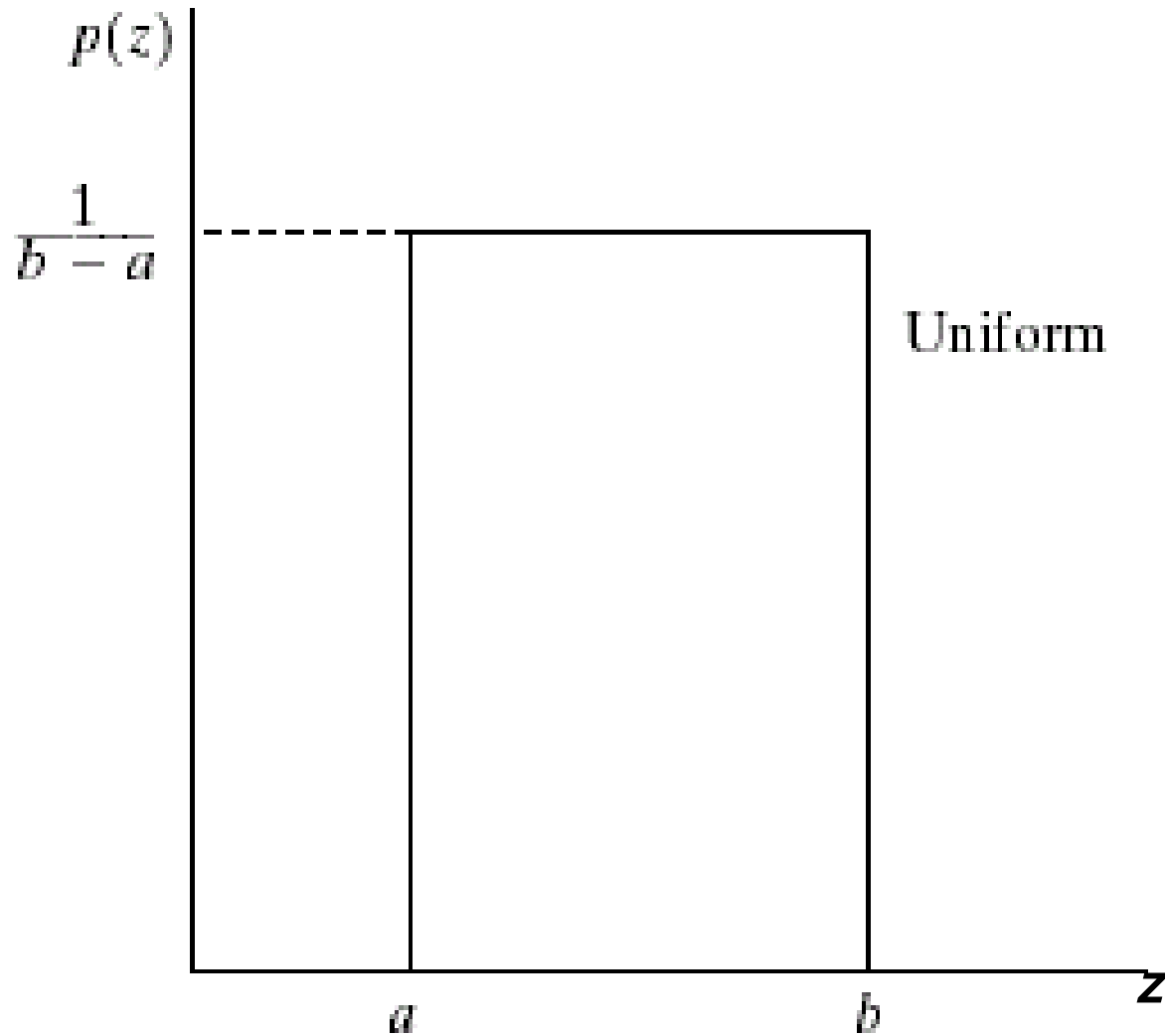
$$p(z) = \begin{cases} \frac{1}{b-a} & \text{if } a \leq z \leq b \\ 0 & \text{otherwise} \end{cases}$$

$$\text{Mean: } \mu = \frac{a+b}{2}$$

$$\text{Variance: } \sigma^2 = \frac{(b-a)^2}{12}$$



# Uniform PDF



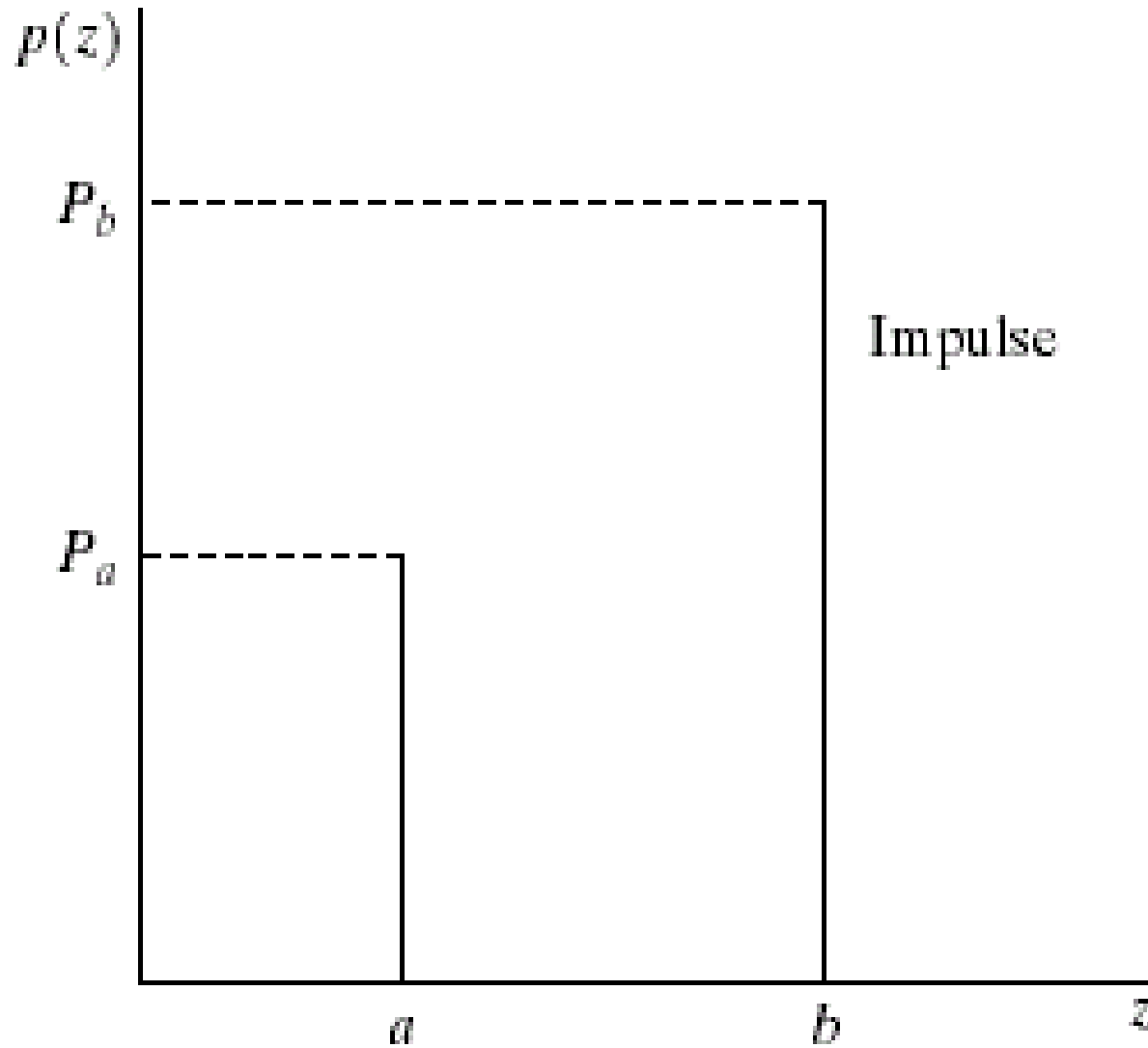
# Impulse (salt-and-pepper) noise

- Quick transients, such as faulty switching during *imaging*

$$p(z) = \begin{cases} P_a & \text{for } z = a \\ P_b & \text{for } z = b \\ 0 & \text{otherwise} \end{cases}$$

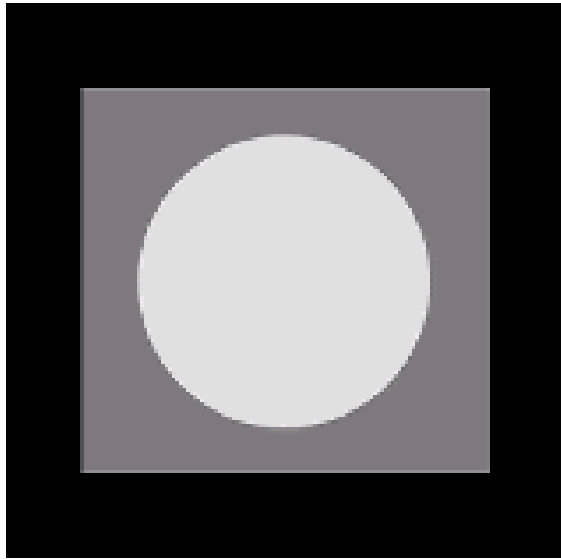
If either  $P_a$  or  $P_b$  is zero, it is called *unipolar*.  
Otherwise, it is called *bipolar*.

# Impulse (salt-and-pepper) noise PDF



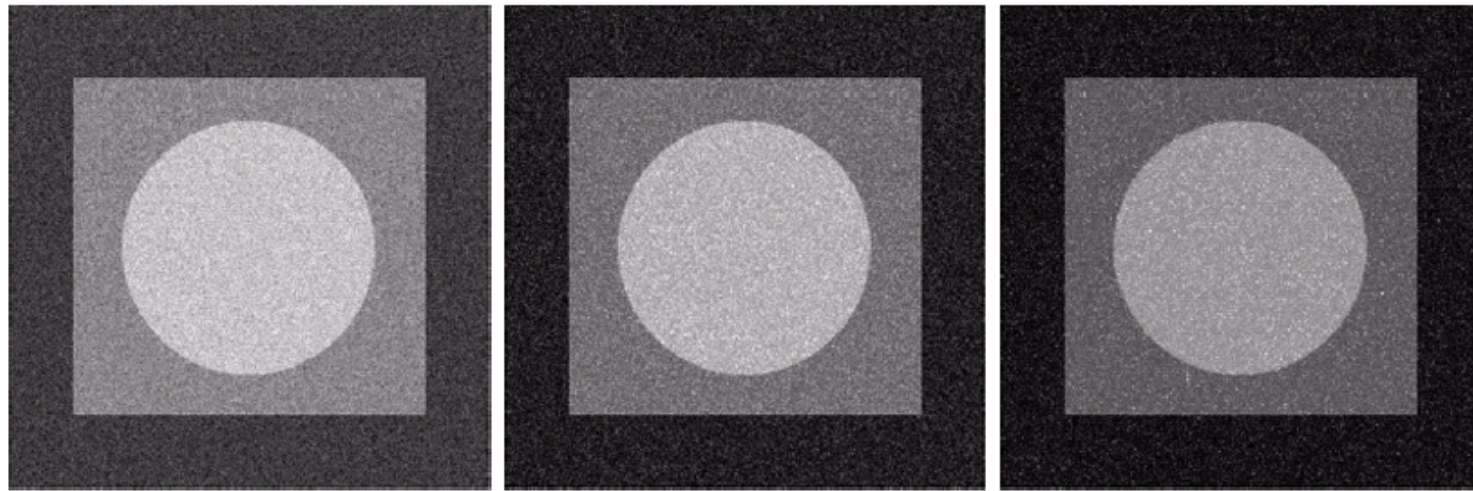
# Image Degradation with Additive Noise

$$g(x, y) = f(x, y) + \eta(x, y)$$

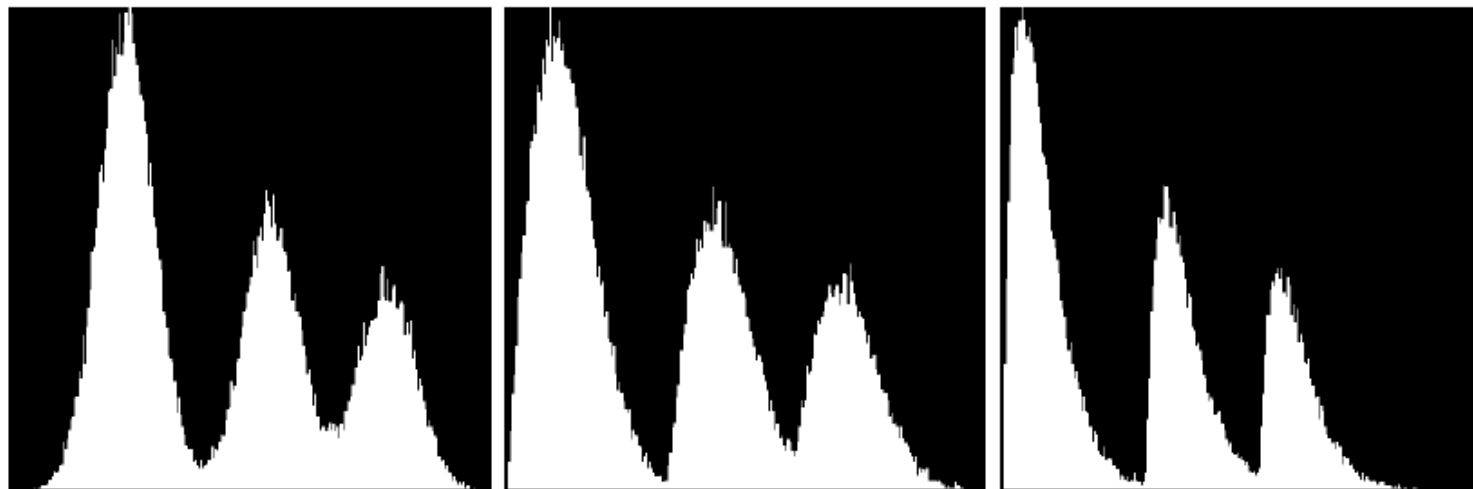
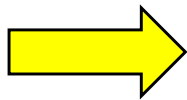


Original image

Degraded images



Histogram

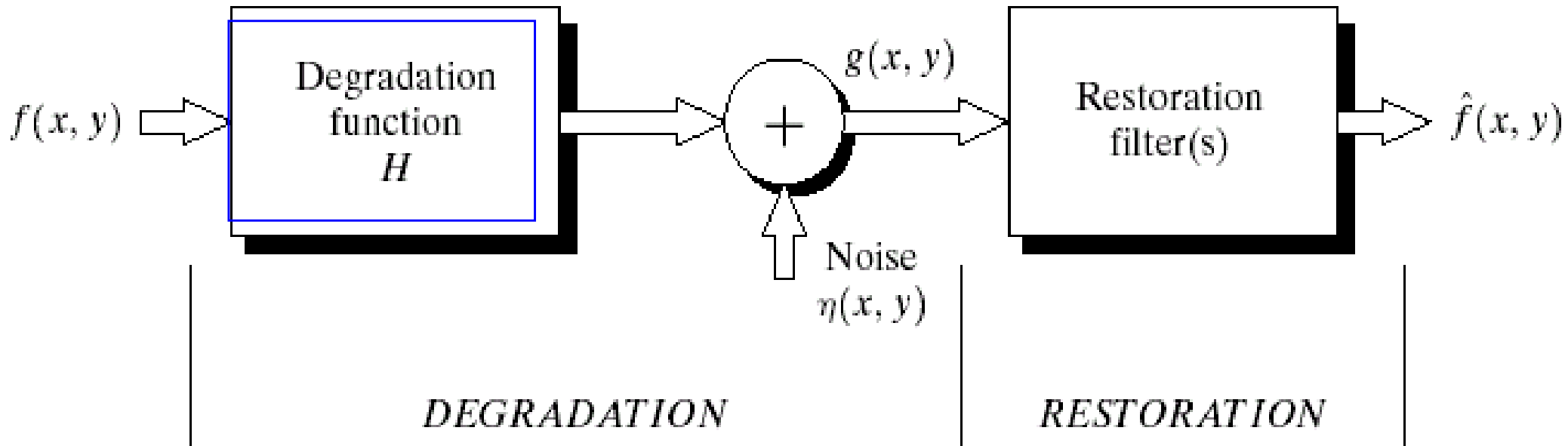


Gaussian

Rayleigh

Gamma

# A model of the image degradation / restoration process



$$\left\{ \begin{array}{l} g(x,y)=f(x,y)*h(x,y)+\eta(x,y) \\ G(u,v)=F(u,v)H(u,v)+N(u,v) \end{array} \right.$$

# Linear, position-invariant degradation

Properties of the degradation function  $H$

- **Linear system**

- $H[af_1(x,y)+bf_2(x,y)]=aH[f_1(x,y)]+bH[f_2(x,y)]$

- **Position(space)-invariant system**

- $H[f(x,y)]=g(x,y)$  is position invariant if

- $H[f(x-\alpha, y-\beta)]=g(x-\alpha, y-\beta)$



# Estimation of Degradation Function

## Degradation model:

$$g(x, y) = f(x, y) * h(x, y) + \eta(x, y)$$

or

$$G(u, v) = F(u, v)H(u, v) + N(u, v)$$

If we know exactly  $h(x, y)$ , regardless of noise, we can do deconvolution to get  $f(x, y)$  back from  $g(x, y)$ .

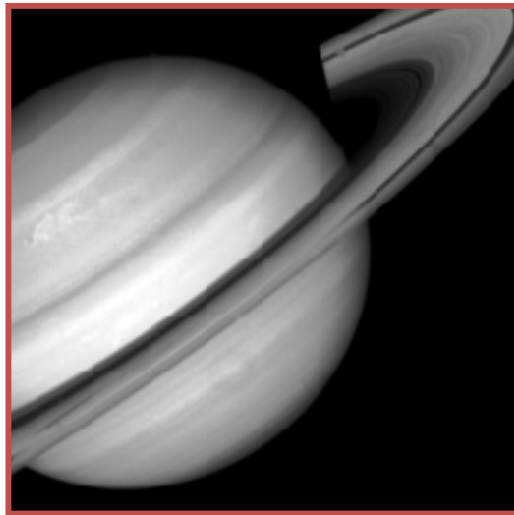
## Methods:

1. Estimation by Image Observation
2. Estimation by Experiment
3. Estimation by Modeling



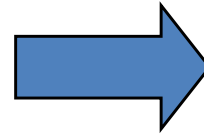
# Estimation by Image Observation

Original image (unknown)

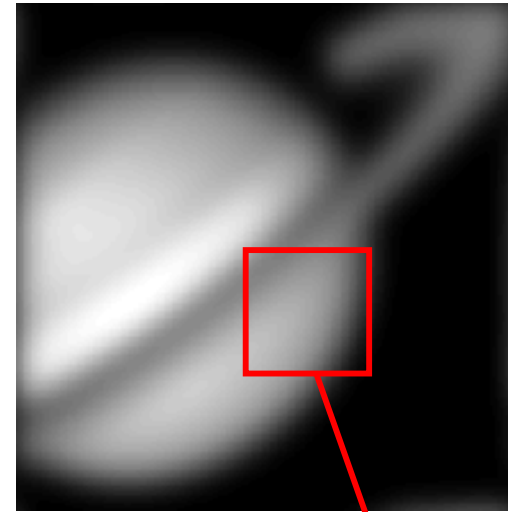


$f(x, y)$

$f(x, y) * h(x, y)$



Degraded image



$g(x, y)$

Observation

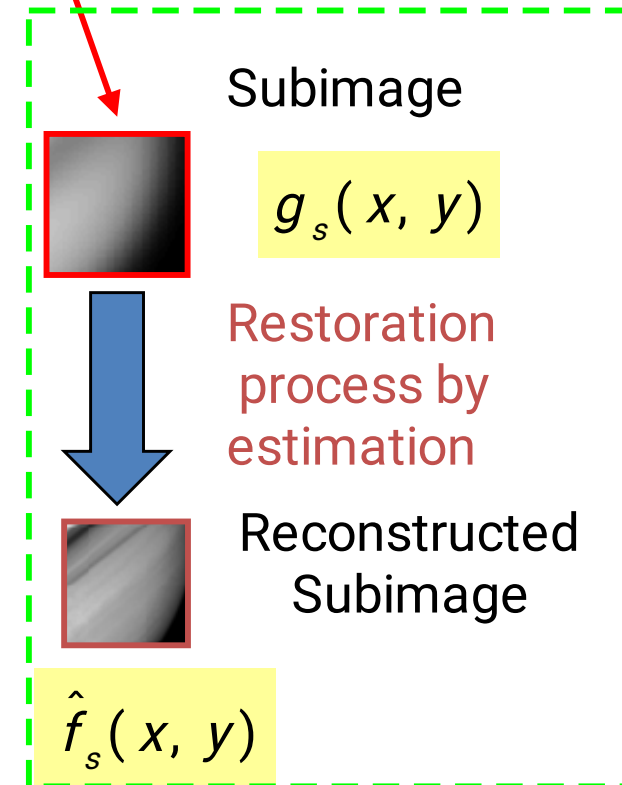
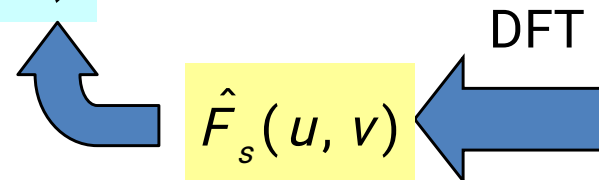
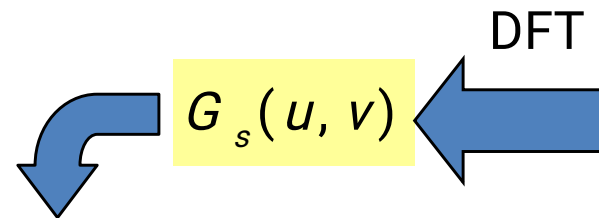


Estimated Transfer function

$$H(u, v) \approx H_s(u, v) = \frac{G_s(u, v)}{\hat{F}_s(u, v)}$$

This case is used when we know only  $g(x, y)$  and cannot repeat the experiment!

3/21/2020

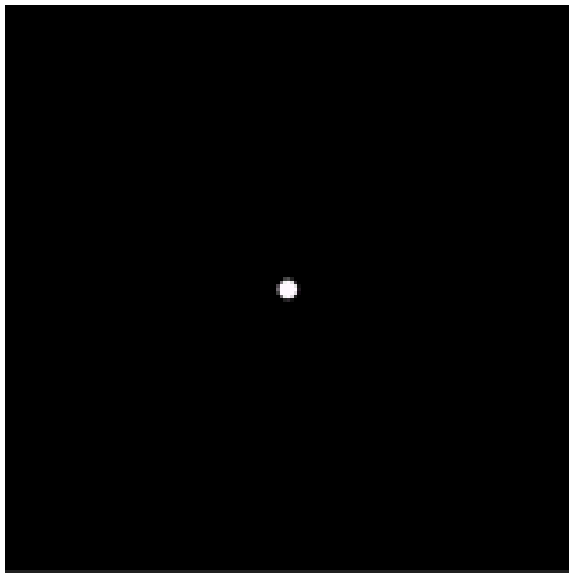




# Estimation by Experiment

Used when we have the same equipment set up

Input impulse image



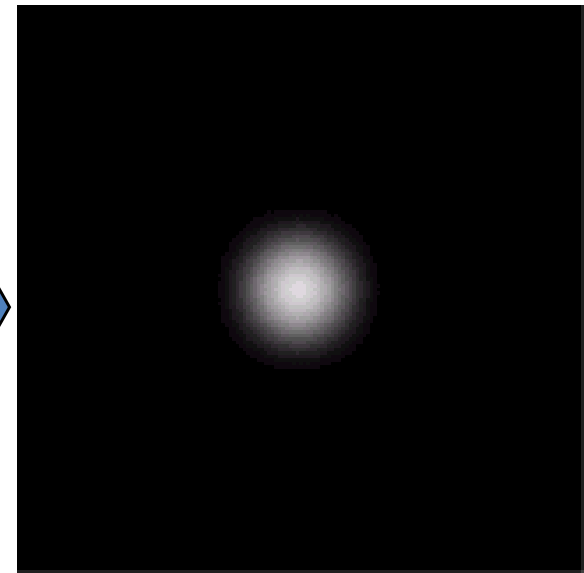
$$A \delta(x, y)$$

DFT

$$\text{DFT} \{A \delta(x, y)\} = A$$

System  
 $H()$

Response image from  
the system



$$g(x, y)$$

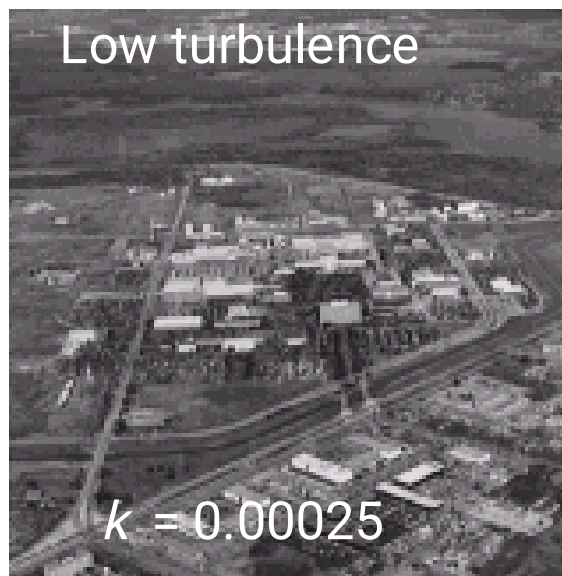
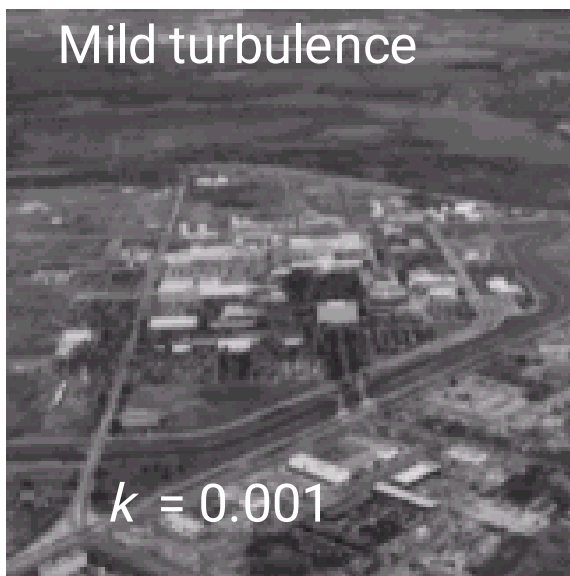
DFT

$$G(u, v)$$

$$H(u, v) = \frac{G(u, v)}{A}$$

# Estimation by Modeling

Used when we know physical mechanism underlying the image formation process that can be expressed mathematically.



Example:

Atmospheric  
Turbulence model

$$H(u, v) = e^{-k(u^2 + v^2)^{5/6}}$$



# Estimation by Modeling: Motion Blurring

Assume that camera velocity is

$$(x_0(t), y_0(t))$$

The blurred image is obtained by

$$g(x, y) = \int_0^T f(x + x_0(t), y + y_0(t)) dt$$

where  $T$  = exposure time.

$$\begin{aligned} G(u, v) &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} g(x, y) e^{-j2\pi(ux+vy)} dx dy \\ &= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} \left[ \int_0^T f(x + x_0(t), y + y_0(t)) dt \right] e^{-j2\pi(ux+vy)} dx dy \\ &= \int_0^T \left[ \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(x + x_0(t), y + y_0(t)) e^{-j2\pi(ux+vy)} dx dy \right] dt \end{aligned}$$



# Estimation by Modeling: Motion Blurring (cont.)

$$\begin{aligned} G(u, v) &= \int_0^T \left[ \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} f(x + x_0(t), y + y_0(t)) e^{-j2\pi(ux + vy)} dx dy \right] dt \\ &= \int_0^T [F(u, v) e^{-j2\pi(ux_0(t) + vy_0(t))}] dt \\ &= F(u, v) \int_0^T e^{-j2\pi(ux_0(t) + vy_0(t))} dt \end{aligned}$$

Then we get, the motion blurring transfer function:

$$H(u, v) = \int_0^T e^{-j2\pi(ux_0(t) + vy_0(t))} dt$$



For constant motion

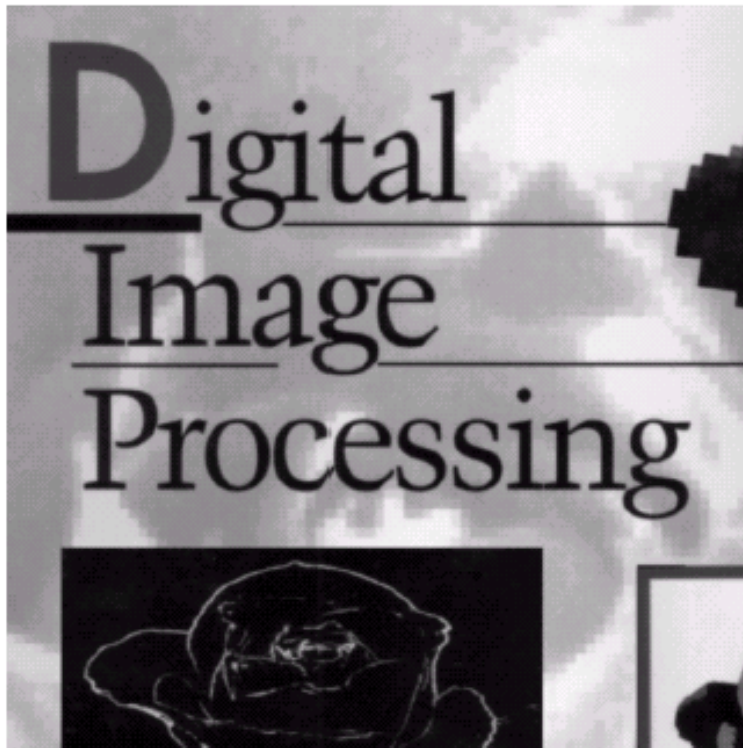
$$(x_0(t), y_0(t)) = (at, bt)$$

$$H(u, v) = \int_0^T e^{-j2\pi(ua + vb)t} dt = \frac{T}{\pi(ua + vb)} \sin(\pi(ua + vb)) e^{-j\pi(ua + vb)}$$

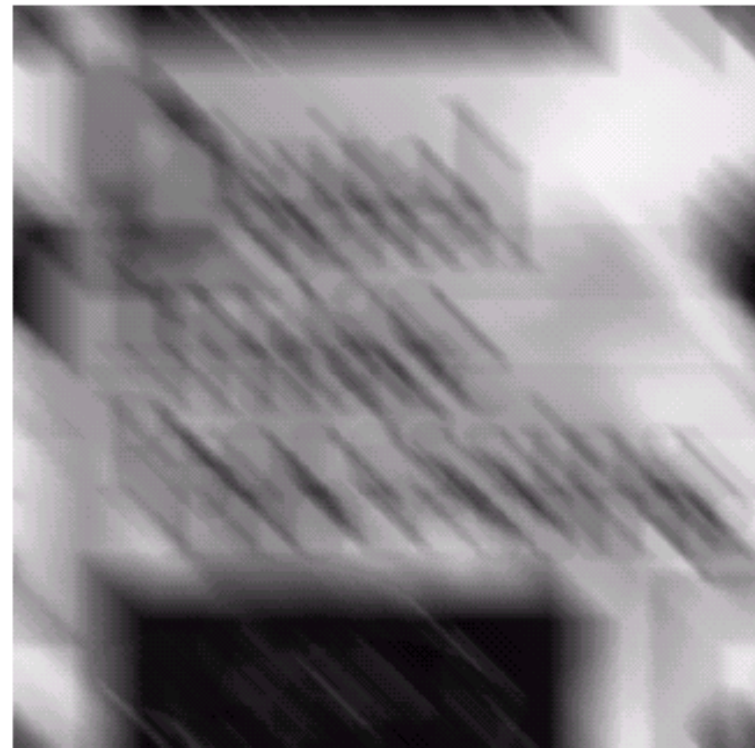
# Motion Blurring Example

For constant motion

$$H(u, v) = \frac{T}{\pi(ua + vb)} \sin(\pi(ua + vb)) e^{-j\pi(ua + vb)}$$



Original image



Motion blurred image  
 $a = b = 0.1, T = 1$

# Inverse Filtering

From degradation model:

$$G(u, v) = F(u, v)H(u, v) + N(u, v)$$

after we obtain  $H(u, v)$ , we can estimate  $F(u, v)$  by the inverse filter:

$$\hat{F}(u, v) = \frac{G(u, v)}{H(u, v)} = F(u, v) + \frac{N(u, v)}{H(u, v)}$$



Noise is enhanced  
when  $H(u, v)$  is small.

To avoid the side effect of enhancing noise, we can apply this formulation to freq. component  $(u, v)$  with in a radius  $D_0$  from the center of  $H(u, v)$ .

In practical, the inverse filter is not popularly used.

# Inverse Filtering Contd...

- Divide equation one by  $H(u,v)$

- $$\frac{G(u,v)}{H(u,v)} = \frac{F(u,v)H(u,v)+N(u,v)}{H(u,v)} \dots\dots(2)$$

- We know that  $\hat{F}(u,v) = \frac{G(u,v)}{H(u,v)}$

- Substitute  $\hat{F}(u,v)$  in eqn (2)

- $$\hat{F}(u,v) = F(u,v) + \frac{N(u,v)}{H(u,v)}$$

- If noise is zero the estimated image  $\hat{F}(u,v)$  is equal to original image, but noise will not be properly removed in inverse filtering.

# Inverse Filtering

- **Limitations:**

1. Even if the degradation function is known the undegraded image cannot be recovered exactly because  $N(u,v)$  is the random function which is not known.
2. If the degradation function has '0' or small value the ratio  $\frac{N(u,v)}{H(u,v)}$  easily dominates the estimate  $F(u,v)$  one approach to get ride of 0 (or) small value problem to limits the filter frequency to the value near the origin.



# WIENER FILTERING

- Inverse filtering has no explicit provision for handling noise but the wiener filtering it incorporates both degradation function, statistical characteristics of noise taken into the restoration process.
  - $e^2 = E[(f-\hat{f})^2]$
- Objective of the wiener filter is to find the estimate of uncorrupted image  $f$ , such that the mean square error is minimize the wiener filter is optimum filter
- **Diagram**



# Wiener Filtering Contd...

- The error between the input signal and the estimated signal is given by the mean square error.
  - $e(x,y) = f(x,y) - \hat{f}(x,y)$
  - $E[f(x,y) - \hat{f}(x,y)] = 0$
- According to the principle of orthogonality the expected value of  $f(x,y) - \hat{f}(x,y)$  totally orthogonal with  $g(x,y)$  is zero.

$$E[f(x,y) - \hat{f}(x,y)g(x,y)] = 0$$

$$\hat{f}(x,y) = g(x,y)*r(x,y)$$

$$E[f(x,y) - (g(x,y)*r(x,y))g(x,y)] = 0$$

$$E[f(x,y)g(x,y)] = E[(r(x,y)*g(x,y))g(x,y)]$$

$$= E\left\{\left[\sum_{k=-\infty}^{\infty} \sum_{l=-\infty}^{\infty} r(x-k, y-l)g(k,l)\right]g(x,y)\right\}$$



# Wiener Filtering Contd...

- $S_{gg}(u,v) = H(u,v)H^*(u,v) \cdot S_{ff}(u,v)$   
 $= |H(u,v)|^2 S_{ff}(u,v)$

$$R(u,v) = \frac{S_{fg}(u,v)}{S_{gg}(u,v)} = \frac{H^*(u,v)S_{ff}(u,v)}{|H(u,v)|^2 S_{ff}(u,v) + S_{\eta}(u,v)}$$

With presence of noise  $S_{gg}(u,v) = |H(u,v)|^2 S_{ff}(u,v) + N(u,v)$

$$S_{\eta}(u,v) = |N(u,v)|^2$$

$$R(u,v) = \frac{H^*(u,v)S_{ff}(u,v)}{|H(u,v)|^2 S_{ff}(u,v) + S_{\eta}(u,v)} = \frac{\hat{F}(u,v)}{G(u,v)}$$

$$\hat{F}(u,v) = R(u,v)G(u,v)$$

Multiply and divide by  $H(u,v)$  in  $R(u,v)$  and sub in  $\hat{F}(u,v)$

$$\hat{F}(u,v) = \left[ \frac{1}{H(u,v)} \frac{[H^*(u,v)H(u,v)]S_{ff}(u,v)}{[|H(u,v)|^2 S_{ff}(u,v) + S_{\eta}(u,v)]} \right] G(u,v)$$



# Wiener Filtering Contd...

- $\hat{F}(u, v) = \left[ \frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + \frac{S_{\eta}(u, v)}{S_{ff}(u, v)}} \right] G(u, v)$
- Wiener filter also known as minimum mean square filter or least mean square filter.
- Wiener filter does not have the same problem as the inverse filter unless both  $H(u, v)$  and  $S_{\eta}(u, v)$  are zero for the same value of  $u$  &  $v$
- $H(u, v)$  = degradation function
- $H^*(u, v)$  = complex conjugate of  $H(u, v)$
- $|H(u, v)|^2 = H^*(u, v) H(u, v)$
- $S_{\eta}(u, v) = |N(u, v)|^2 =$  Power spectrum of the noise
- $S_f(u, v) = |F(u, v)|^2 =$  Power spectrum of an undegraded image.



# Wiener Filtering

- Consideration:

1. When a noise is zero

$$\eta(x,y)=0, S\eta(u,v) =0$$

$$\hat{F}(u, v) = \frac{G(u,v)}{H(u,v)}$$

It reduces to inverse filtering



2. IF  $H(u,v)=1$

$$\hat{F}(u, v) = \left[ \frac{G(u,v)S_{ff}(u,v)}{S_{ff}(u,v) + S\eta(u,v)} \right]$$

$$\frac{G(u,v) \frac{S_{ff}(u,v)}{S\eta(u,v)}}{\frac{S_{ff}(u,v)}{S\eta(u,v)} + 1}$$

# Wiener Filtering

- Signal to Noise ratio  $\frac{S_{ff}(u,v)}{S_{\eta}(u,v)}$

3. Signal to noise ratio is greater than 1



$$\frac{S_{ff}(u,v)}{S_{\eta}(u,v)} \gg 1$$

Then  $\hat{F}(u, v) = G(u,v)$  --- Here the wiener filter act as a all pass filters.

## ADVANTAGES:

1. The wiener filter does not have zero value problem until both  $H(u,v)$  and  $S_{\eta}(u, v)$  is equal to zero.
2. The result obtained by wiener filter is more closer to the original image than inverse filter.

# Approximation of Wiener Filter

Wiener Filter Formula:

$$\hat{F}(u, v) = \left[ \frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + S_{\eta}(u, v) / S_f(u, v)} \right] G(u, v)$$

Difficult to estimate

Approximated Formula:

$$\hat{F}(u, v) = \left[ \frac{1}{H(u, v)} \frac{|H(u, v)|^2}{|H(u, v)|^2 + K} \right] G(u, v)$$

In Practice,  $K$  is chosen manually to obtain the best visual result!

# Constrained Least Squares Filter

Degradation model:

$$g(x, y) = f(x, y) * h(x, y) + \eta(x, y)$$

Aims to find the minimum of a criterion function

$$C = \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} [\nabla^2 f(x, y)]^2$$

Subject to the constraint

$$\|g - H\hat{f}\|^2 = \|\eta\|^2$$

Constrained least square filter is given by,

$$\hat{F}(u, v) = \left[ \frac{H^*(u, v)}{|H(u, v)|^2 + \gamma |P(u, v)|^2} \right] G(u, v)$$

where

$P(u, v) =$  Fourier transform of  $p(x, y) =$

In matrix form,

$$g = Hf + \eta$$



where

$$\|w\|^2 = w^T w$$

$$\begin{bmatrix} 0 & -1 & 0 \\ -1 & 4 & -1 \\ 0 & -1 & 0 \end{bmatrix}$$

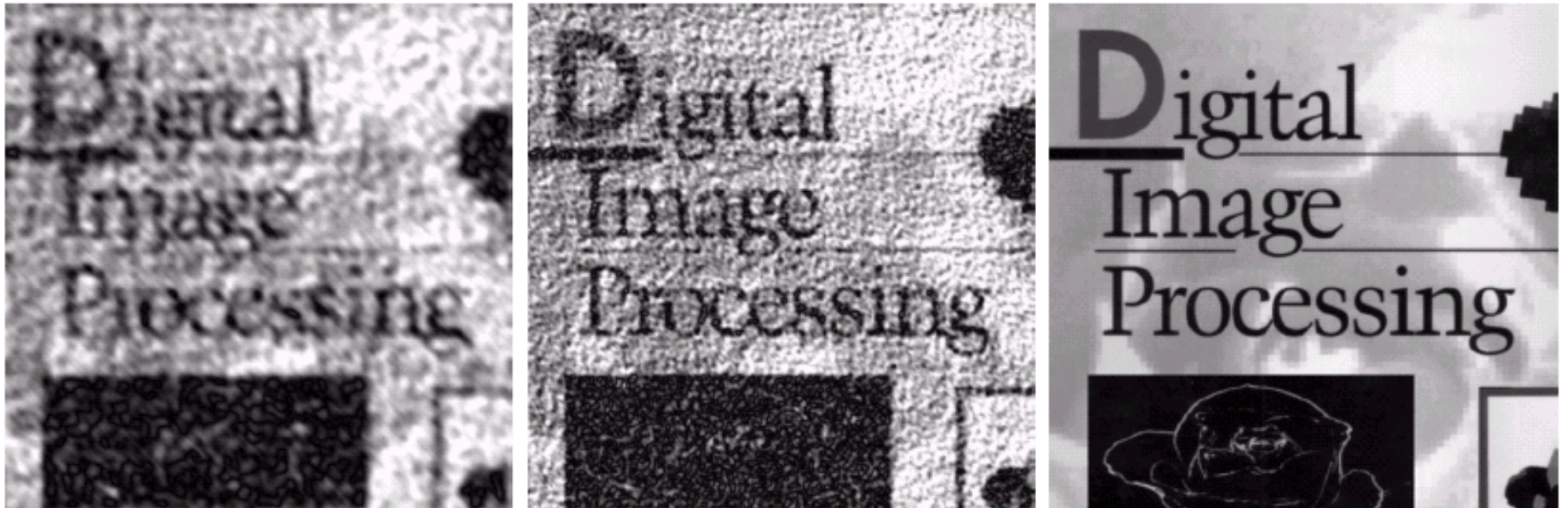


# Constrained Least Squares Filter: Example

Constrained least square filter

$$\hat{F}(u, v) = \left[ \frac{H^*(u, v)}{|H(u, v)|^2 + \gamma |P(u, v)|^2} \right] G(u, v)$$

$\gamma$  is adaptively adjusted to achieve the best result.



Results from the previous slide obtained from the constrained least square filter

# Constrained Least Squares Filter: Adjusting $\gamma$

Define

$$\mathbf{r} = \mathbf{g} - \mathbf{H} \hat{\mathbf{f}}$$

It can be shown that

$$\phi(\gamma) = \mathbf{r}^T \mathbf{r} = \|\mathbf{r}\|^2$$

We want to adjust gamma so that

$$\|\mathbf{r}\|^2 = \|\boldsymbol{\eta}\|^2 \pm a \rightarrow \textcircled{1}$$

where  $a$  = accuracy factor

1. Specify an initial value of  $\gamma$

2. Compute

$$\|\mathbf{r}\|^2$$

3. Stop if is satisfied

Otherwise return step 2 after increasing  $\gamma$  if

1

or decreasing  $\gamma$  if

Use the new value of  $\gamma$  to recompute

$$\|\mathbf{r}\|^2 < \|\boldsymbol{\eta}\|^2 - a$$

$$\|\mathbf{r}\|^2 > \|\boldsymbol{\eta}\|^2 + a$$

$$\hat{F}(u, v) = \left[ \frac{H^*(u, v)}{|H(u, v)|^2 + \gamma |P(u, v)|^2} \right] G(u, v)$$



# Constrained Least Squares Filter: Adjusting $\gamma$ (cont.)

$$\hat{F}(u, v) = \left[ \frac{H^*(u, v)}{|H(u, v)|^2 + \gamma |P(u, v)|^2} \right] G(u, v)$$

$$R(u, v) = G(u, v) - H(u, v) \hat{F}(u, v)$$

$$\|\mathbf{r}\|^2 = \frac{1}{MN} \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} r^2(x, y)$$

For computing

$$\|\mathbf{r}\|^2$$

$$m_\eta = \frac{1}{MN} \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} \eta(x, y)$$

$$\sigma_\eta^2 = \frac{1}{MN} \sum_{x=0}^{M-1} \sum_{y=0}^{N-1} [\eta(x, y) - m_\eta]^2$$

$$\|\mathbf{n}\|^2 = MN [\sigma_\eta^2 - m_\eta]$$

For computing

$$\|\mathbf{n}\|^2$$

# Geometric Transformation

These transformations are often called **rubber-sheet transformations**:  
Printing an image on a rubber sheet and then stretch this sheet according to some predefined set of rules.



A geometric transformation consists of 2 basic operations:

**1. A spatial transformation :**

Define how pixels are to be rearranged in the spatially transformed image.

**2. Gray level interpolation :**

Assign gray level values to pixels in the spatially transformed image.

# Geometric Transformation: Algorithm

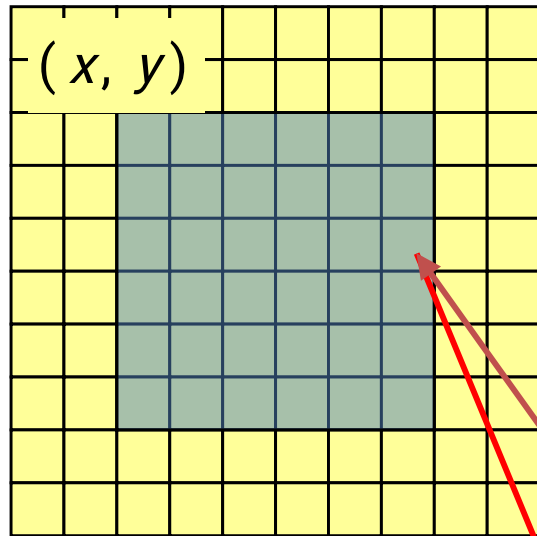
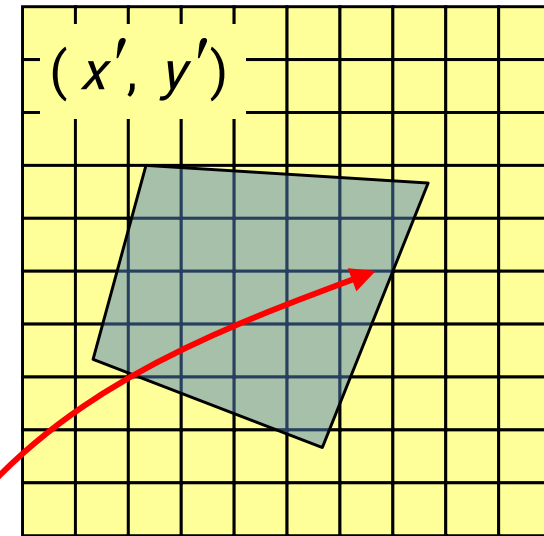


Image  $f$  to be restored



Distorted image  $g$

1

3



1. Select coordinate  $(x, y)$  in  $f$  to be restored

2. Compute

$$x' = r(x, y)$$

$$y' = s(x, y)$$

3. Go to pixel  $(x', y')$  in a distorted image  $g$

4. get pixel value at  $g(x', y')$   
By gray level interpolation,

5. store that value in pixel  $f(x, y)$

5

# Spatial Transformation

To map between pixel coordinate  $(x, y)$  of  $f$  and pixel coordinate  $(x', y')$  of  $g$

$$x' = r(x, y)$$

$$y' = s(x, y)$$

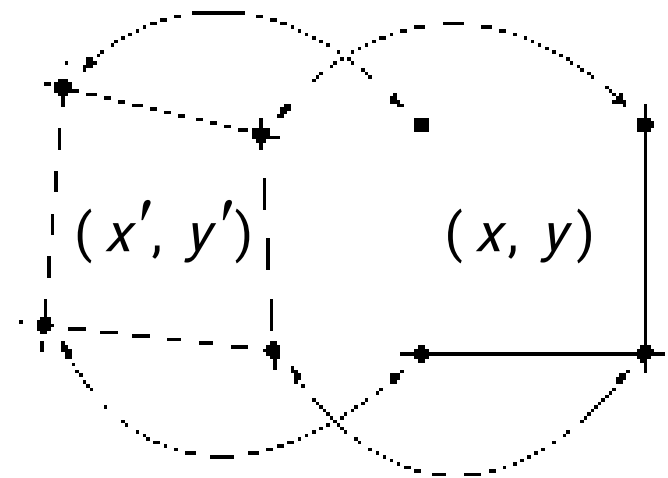
For a bilinear transformation mapping between a pair of Quadrilateral regions

$$x' = r(x, y) = c_1 x + c_2 y + c_3 xy + c_4$$

$$y' = s(x, y) = c_5 x + c_6 y + c_7 xy + c_8$$

To obtain  $r(x, y)$  and  $s(x, y)$ , we need to know 4 pairs of coordinates and its corresponding  $(x, y)$  are called **tiepoints**.

$$(x', y')$$



# Gray Level Interpolation: Nearest Neighbor

Since  $(x', y')$  is not at an integer coordinate, we need to interpolate the value of  $g(x', y')$

Example interpolation methods that can be used:

1. Nearest neighbor selection
2. Bilinear interpolation
3. Bicubic interpolation

